



National Aeronautics and  
Space Administration  
Cosmic Origins  
Program Office



# Cosmic Origins Program Update

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Some Title, Cosmic Origins Program Office

COPAG Workshop

January 6, 2013

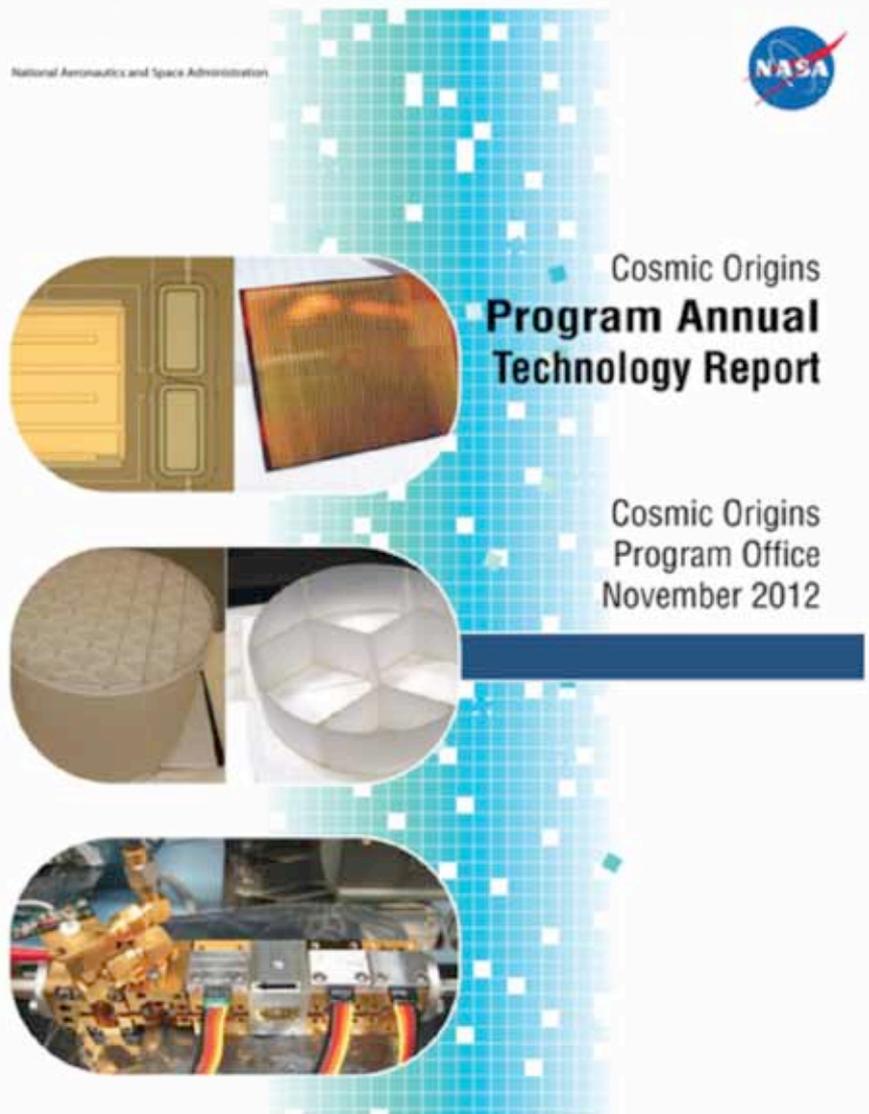
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- **Technology Priorities**
- SAT Proposal
- Funding Outlook

- PCOS and COR each convene a Technology Management Board (TMB) to review community-identified technologies that point toward future missions
- Output is prioritized list of investments; one purpose is as input to SAT call for proposals

- Program Office implements technology maturation, inheriting needs from the broadest community
- Stakeholders identify science needs
- PAG = formal introduction mechanism

1. High QE, large format UV detectors
2. Photon counting UV large-format detectors
3. High Reflectivity UV coatings
4. Large, low-cost, lightweight precision mirrors for UV/visible Telescopes
5. Deployable lightweight precision mirrors for UV/visible Telescopes
6. Very large format, low noise visible/IR detector arrays
7. Photon counting visible/IR detector arrays
8. Large format, low noise Far-IR direct detectors
9. Ultralow-noise far-IR direct detectors
10. Large, cryogenic far-IR telescopes
11. Interferometry for far-IR telescopes
12. High-Performance Sub-Kelvin Coolers
13. Coherent far-IR detector arrays
14. High-Efficiency Cryocoolers
15. High-Efficiency Spectrometers



- Contains status of COR-funded technology developments
- Provides new technology needs prioritization.
- Available online!  
<http://cor.gsfc.nasa.gov>

- The PATR is an annual report that summarizes the Program’s technology development activities for the year. It consists of an Executive Summary and 4 (significant) Sections:
  1. Provides an overview of the Program and its technology development activities
  2. Provides a status of the Program’s technology development activities for the prior year and announces the new SAT award selections.
  3. Summarizes the technology needs obtained from the community.
  4. Provides a prioritized list of technology needs for the coming year to inform the SAT proposal calls and selection decision (information is also used to inform other NASA technology programs such as SBIR and other OCT planning)
- PATR is updated annually and is released in October to support annual technology development planning.

- A Program technology needs identification and prioritization process has been implemented for PCOS and COR since last year
- The objectives of this process are to:
  - Identify technology needs that are applicable and relevant to Program science objectives
  - Then prioritize these needs with respect to scientific priorities, benefits and impacts, timeliness, and effectiveness of investment.
- The outcome of this process is used to:
  - Inform the Program’s call for SAT proposals and other technology development Program planning (SBIR and other OCT activities)
  - Inform technology developers of the Program needs
  - Guide the selection of technology awards to be aligned with Program goals and science objectives
- This process is designed to:
  - Improve the transparency and relevance of Program technology investments
  - Inform the community about and engage it in our technology development process
  - Leverage the technology investments of external organizations by defining a need and a customer



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# Overview of the Technology Needs Prioritization Process



- The community identifies technology needs each summer by working with the PAG or through direct individual submission to the Program Office's website.
- The Program Technology Management Board (TMB) reviews and vets community identified technology needs, defines their priorities, and recommends investment consideration.
  - TMB membership includes senior members of the Program at NASA HQ and in the Program Office, and when needed, technical expert(s) from the community.
- The TMB prioritizes the technology needs based on a published set of criteria that includes an 11-point assessment that addresses scientific priorities (Decadal Survey), benefits and impacts, timeliness, and effectiveness of investment.
- The technology needs and the resulting priorities are published each year in the PATR.

## COR Technology Needs Prioritization

High QE, large format UV detectors	<p><b>Priority 1.</b> Contains technology needs that the TMB has determined to be of the highest interest to the Cosmic Origins program and recommends that they <b>should</b> be invested in first, when funding is available</p>
Photon counting UV large-format detectors	
High Reflectivity UV coatings	
Ultralow-noise Far-IR direct detectors	
Very large format, low noise Optical/IR detector arrays	<p><b>Priority 2:</b> Contains technology needs that the TMB feels are worthy of pursuit and <b>would</b> be invested in, if funding allows</p>
Large, low-cost, light-weight precision mirrors for Ultra-Stable Large Aperture UV/Optical Telescopes	
Large format, low noise Far-IR direct detectors	
Photon counting Optical/IR detector arrays	
Heterodyne Far-IR receiver arrays	
High efficiency cryocoolers	<p><b>Priority 3:</b> Contains technology needs that are deemed to be supportive of COR objectives but, for various reasons, do not warrant investment at the present, although they <b>could</b> be invested in, if significant additional funding is available</p>
High efficiency UV multi-object spectrometers	
Large, cryogenic far-IR telescopes	
High Performance Sub-Kelvin Coolers	
Deployable light-weight precision mirrors for future Very Large Aperture UV/Optical Telescopes	
Interferometry for far-IR telescopes	

- The PATR is published annually to inform the community of the Program's technology development activities, technology needs and their prioritization.
- The community identifies technology needs by working with the PAG or through direct submission using the program web site.
- The Program TMB prioritizes these needs based on a published set of criteria
- These technology needs and priorities are published each year in the PATR along with the status of technologies that were funded the previous year.
- The program references these priorities and this report over the following year as the calls for technology proposals are drafted and investment decisions are made.
- This process improves the transparency and relevance of technology investments, provides the community a voice in the process, and leverages the technology investments of external organizations by defining needs and a customer.

- Technology Priorities
- **SAT Proposal**
- Funding Outlook

- High-QE, large-format UV detectors – QE (>70%), large-format (>2k × 2k) detectors for operation at 100–400 nm or broader .
- Photon-counting, large-format UV detectors – For spectroscopy, high QE (>50%), very low-noise (<10<sup>-7</sup> ct/pixel/s), large-format (>2k × 2k) photon-counting detectors for operation at 100–400 nm or broader .
- UV coatings –high reflectivity, high uniformity, and wide bandpasses, operating from visible to wavelengths below 100 nm .
- Ultra-low-noise far-IR direct detectors – For spectroscopy at wavelengths between ~30 μm and ~300 μm;  
NEP ≈ 3×10<sup>-21</sup> W/√Hz arrayable in a close-packed configuration in at least one direction

- Detectors:
  - High-QE, large format, photon counting & ultralow-noise
  - From Extreme UV to Far-IR
  - Unchanged?

- UV Coatings:
  - Reflective, antireflective, wavelength-selective
  - Emphasis on Lyman UV (900-1300Å)
  - Changes from previous:
    - Allows A/R coatings; dichroics, filters
    - De-emphasizes Al + protective layer

- Optics:
  - Scalable manufacturing, test, control to sizes of at least 4m
  - Advances in: areal density, production time, cost, precision, short UV wavelengths, surface controls
  - Changes from previous:
    - Size from 2m  $\rightarrow$  4m
    - Removed cryogenic optics; emphasize Lyman UV

- Added sentence about lower-tier priorities:
  - Heterodyne far-IR receiver arrays for, e.g., SOFIA
  - High performance cryocoolers for space flight for future far-IR mission
- Added language allowing piggyback on suborbital flight (balloon or sounding rocket)
- Added language about reporting requirements for funded SAT PIs

- Technology Priorities
- SAT Proposal
- **Funding Outlook**

# Astrophysics Funding COR

	FY04 Final	FY05 Final	FY06 Final	FY07 Final	FY08 Final	FY09 Final	FY10 Final	FY11 Final	FY12
	\$k	\$k	\$k	\$k	\$k	\$k	\$k	\$k	Projected
Particle Astro	\$ 8,248	\$ 7,671	\$ 8,544	\$ 7,631	\$ 6,672	\$ 8,201	\$ 8,260	\$ 8,243	\$ 8,585
High Energy	\$ 14,548	\$ 13,693	\$ 14,779	\$ 12,782	\$ 12,406	\$ 13,886	\$ 14,110	\$ 13,911	\$ 14,548
UV/Opt/IR/ Sub-mm	\$ 20,409	\$ 18,742	\$ 21,851	\$ 17,442	\$ 19,094	\$ 22,353	\$ 21,534	\$ 21,295	\$ 23,032
Other	\$ 1,019	\$ 854	\$ 338	\$ 394	\$ 594	\$ 670	\$ 673	\$ 641	\$ 1,627
<b>APRA Total</b>	<b>\$ 44,224</b>	<b>\$ 40,960</b>	<b>\$ 45,511</b>	<b>\$ 38,250</b>	<b>\$ 38,765</b>	<b>\$ 45,110</b>	<b>\$ 44,577</b>	<b>\$ 44,090</b>	<b>\$ 47,791</b>
Orig Solar Systems	\$ 4,209	\$ 3,872	\$ 4,150	\$ 3,673	\$ 2,965	\$ 3,000	\$ 2,807	\$ 2,944	\$ 2,978
Astro Theory Program	\$ 7,860	\$ 7,363	\$ 10,245	\$ 10,227	\$ 11,696	\$ 11,890	\$ 12,262	\$ 12,577	\$ 13,226
<b>R&amp;A (399131)</b>	<b>\$ 56,293</b>	<b>\$ 52,195</b>	<b>\$ 59,906</b>	<b>\$ 52,150</b>	<b>\$ 53,426</b>	<b>\$ 60,000</b>	<b>\$ 59,646</b>	<b>\$ 59,611</b>	<b>\$ 63,995</b>
ADAP/LTSA	\$ 16,986	\$ 15,700	\$ 15,189	\$ 12,641	\$ 12,013	\$ 14,384	\$ 13,258	\$ 14,132	\$ 16,320
<b>Core Research</b>	<b>\$ 73,279</b>	<b>\$ 67,895</b>	<b>\$ 75,095</b>	<b>\$ 64,791</b>	<b>\$ 65,439</b>	<b>\$ 74,384</b>	<b>\$ 72,904</b>	<b>\$ 73,743</b>	<b>\$ 80,315</b>
TPF/FS Beyond Einstein FS	\$ 2,000	\$ 2,000		(Foundation Science; now in ATP)					
ASMCS (399131)					\$ 3,452	\$ 442			
PCOS SR&T							\$ 968	\$ 184	
Technology Fellows									\$ 600
<b>TOTAL</b>	<b>\$ 79.3M</b>	<b>\$ 72.9M</b>	<b>\$ 77.1M</b>	<b>\$ 64.8M</b>	<b>\$ 68.9M</b>	<b>\$ 74.8M</b>	<b>\$ 73.9M</b>	<b>\$ 73.9M</b>	<b>\$ 80.9M</b>
		\$7M cut	smaller cut	15% cut	partial recovery	more recovery	flat	flat	growth!

In response to the Astro2010 Decadal Survey recommendations:

- The budget for research awards increased by 10% in FY12
- Theory and Computation Networks: AAAC studying NASA-NSF program
- Suborbital program (payloads, balloons) growth deferred

From Hertz 11/2012



# Future SR&T?



	FY 11	FY 12	FY 13	FY 14	FY 15	FY 16	FY 17
				<i>(FY14-17 estimates are notional)</i>			
<b>Astrophysics</b>	<b>631.1</b>	<b>672.7</b>	<b>659.4</b>	<b>703.0</b>	<b>693.7</b>	<b>708.9</b>	<b>710.2</b>
<u>Astrophysics Research</u>	<u>146.9</u>	<u>164.1</u>	<u>176.2</u>	<u>189.1</u>	<u>205.1</u>	<u>211.5</u>	<u>218.7</u>
Astrophysics Research and Analysis	59.6	64.6	64.2	65.5	66.8	68.2	69.5
Balloon Project	26.8	31.6	31.3	31.2	32.8	34.2	34.3
<u>Other Missions and Data Analysis</u>	<u>60.5</u>	<u>67.9</u>	<u>80.6</u>	<u>92.3</u>	<u>105.4</u>	<u>109.2</u>	<u>114.8</u>
Keck Single Aperture	2.2	2.3	2.4	2.4	2.5	2.5	2.5
Astrophysics Data Analysis Program	14.1	16.3	18.3	18.5	18.5	19.1	19.1
Astrophysics Data Curation and Archival	20.8	20.1	20.0	19.6	21.7	22.1	22.2
Astrophysics Senior Review			16.3	24.5	33.5	35.2	40.0
Education and Public Outreach	13.2	15.4	10.1	10.1	10.1	10.1	10.1
Directorate Support - Space Science	10.1	13.7	13.5	13.9	14.0	14.5	14.5
Directed Research and Technology				3.3	5.2	5.6	6.4
<u>Cosmic Origins</u>	<u>229.1</u>	<u>237.3</u>	<u>240.4</u>	<u>228.5</u>	<u>215.1</u>	<u>205.3</u>	<u>205.7</u>
Hubble Space Telescope (HST)	91.7	95.7	98.3	98.3	94.3	90.2	90.5
SOFIA	79.9	84.2	85.5	88.0	88.0	86.0	85.9
<u>Other Missions And Data Analysis</u>	<u>57.6</u>	<u>57.4</u>	<u>56.6</u>	<u>42.2</u>	<u>32.8</u>	<u>29.1</u>	<u>29.3</u>
Spitzer Space Telescope	22.7	17.8	9.8				
Herschel	24.6	24.0	20.8	15.8	5.8		
<b>Cosmic Origins SR&amp;T</b>	<b>7.9</b>	<b>10.6</b>	<b>19.4</b>	<b>19.5</b>	<b>20.7</b>	<b>21.7</b>	<b>21.8</b>
Cosmic Origins Future Missions	0.7	1.0	1.7	1.7	1.0	2.0	2.0
Cosmic Origins Program Management	1.7	4.0	4.9	5.2	5.3	5.4	5.5

*From Hertz 11/2012; notional FY14+*

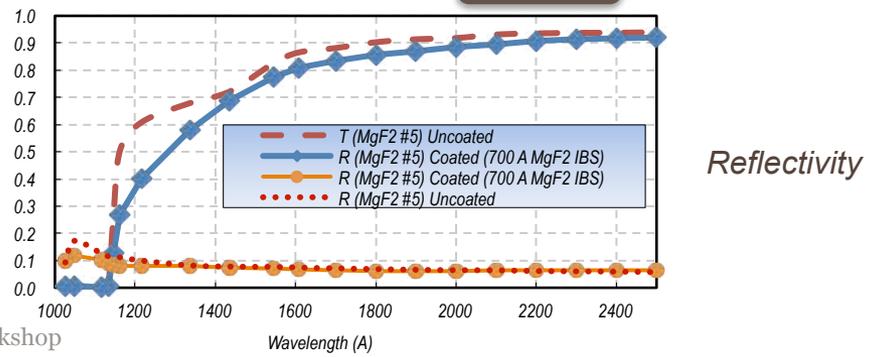
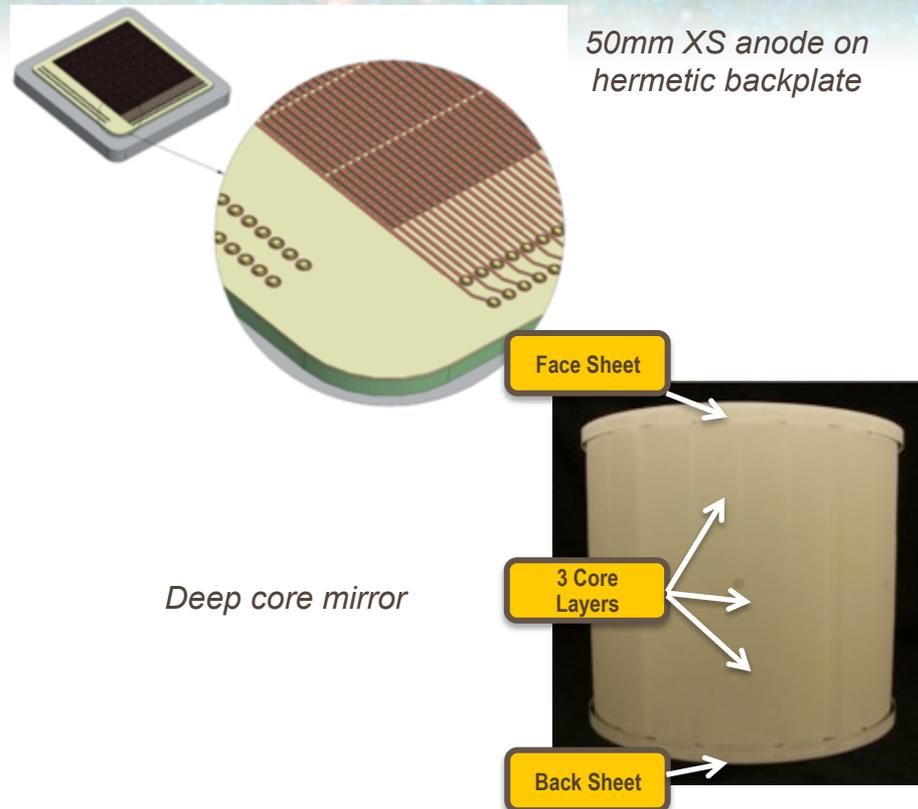
- Strategic Astrophysics Technologies (SAT) – Competed via ROSES
  - Program priorities established and documented in the COR Program Annual Technology Report (PATR)
  - SAT call for proposals is informed by the priorities in PATR
  - Selection of the proposed technologies is based in part on the program priorities – thus consistency is ensured!
- Targeted Technologies (formerly “Directed”)
  - Tied to a specific mission concept
  - Documented in a Technology Development Plan
  - Vetted through a TMB
- Unique Infrastructure – Directed/Competed
  - Capability that serves the community
  - Examples include optical test beds and detector development and characterization labs at NASA Centers or academic institutions

<b>Proposal Title</b>	<b>PI</b>	<b>Institution</b>	<b>Area</b>
<b>Advanced UVOIR Mirror Technology Development for Very Large Space Telescopes</b>	Phillip Stahl	Marshall Space Flight Center	<i>Advanced, Normal Incidence Optics</i>
<b>High performance cross-strip micro-channel plate detector systems for spaceflight experiments</b>	John Vallergera	UC Berkeley	<i>UV Detectors</i>
<b>Enhanced MgF<sub>2</sub> and LiF Over-coated Al Mirrors for FUV Space Astronomy</b>	Manuel Quijada	Goddard Space Flight Center	<i>Ultraviolet Coatings</i>

# COR Technology Progress at End of Year



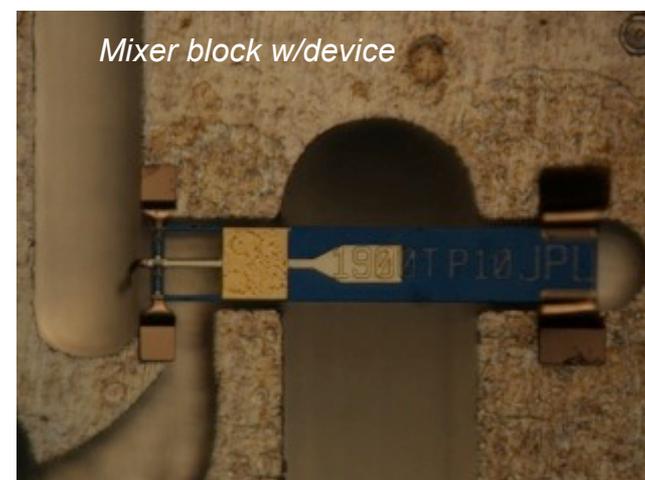
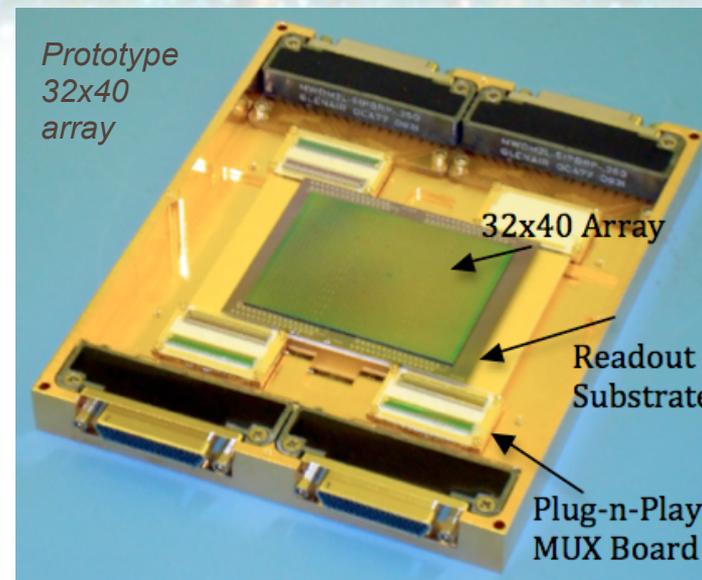
- Cross Strip MCP detector systems for spaceflight – John Vallerga/UC Berkeley
  - Major designs complete & in fabrication
  
- Advanced UVOIR Mirror Technology Development Dr. Phil Stahl/MSFC
  - Demonstrated deep core mirror fab
  
- Enhanced MgF<sub>2</sub> and LiF Over-coated Al Mirrors for FUV Space Astronomy Dr. Manuel Quijada/GSFC
  - Coatings of MgF<sub>2</sub>, LuF<sub>3</sub> made



# COR Technology Progress at End of Year



- Superconducting TES Bolometer Arrays for Far Infrared Imaging, Spectroscopy, and Polarimetry – Harvey Moseley/GSFC
  - All component technologies demonstrated
  - Selected for SOFIA/HAWC+ instrument upgrade (SOFIA 2<sup>nd</sup> Gen); starting now
- Heterodyne Technology for SOFIA – Dr. Paul Goldsmith/JPL
  - Completed mixer blocks and multipliers for 1.9THz



<b>Proposal Title</b>	<b>PI</b>	<b>Institution</b>	<b>Area</b>
<b>Ultraviolet Coatings, Materials and Processes for Advanced Telescope Optics</b>	Kunjithapatham Balasubramanian	JPL	<i>UV Coatings</i>
<b>Kinetic Inductance Detector Imaging Arrays for Far-Infrared Astrophysics</b>	Jonas Zmuidzinas	Caltech	<i>Far-IR Detectors</i>
<b>Improvement of the Performance of Near-Infrared Detectors for NASA Astrophysics Missions</b>	Selmer Anglin	Teledyne	<i>UVOIR Detectors</i>
<b>H4RG Near-IR Detector Array with 10 Micron Pixels for WFIRST</b>	Bernie Rauscher	GSFC	<i>UVOIR Detectors</i>
<b>High Efficiency Detectors in Photon Counting and Large Focal Plane Arrays for Astrophysics Missions</b>	Shouleh Nikzad	JPL	<i>UVOIR Detectors</i>

- Articulate the key drivers for Inflation Probe, SOFIA, & future Far-IR Observatory
  - Science Requirements
  - Technology development needs
  - Timeliness
- Identify alternative mechanisms for achieving PCOS/  
COR science goals
  - Suborbital concepts
  - Coordination with ground-based facilities
  - Technology leading to Explorers, etc.
- Identifying important areas for future SAT calls



# Help?



- Technology Needs – Add to PATR/SAT
- Alternatives to Major Missions
- Things Missing from the Portfolio?